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ASTRONOMY TODAY ^{9e}

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BRIEF CONTENTS

PART I: Astronomy and the Universe 2

- 1 Charting the Heavens: The Foundations of Astronomy 5
- 2 The Copernican Revolution: The Birth of Modern Science 33
- 3 Radiation: Information from the Cosmos 59
- 4 Spectroscopy: The Inner Workings of Atoms 79
- 5 Telescopes: The Tools of Astronomy 99

PART II: Our Planetary System 132

- 6 The Solar System: Comparative Planetology and Formation Models 134
- 7 Earth: Our Home in Space 160
- 8 The Moon and Mercury: Scorched and Battered Worlds 187
- 9 Venus: Earth's Sister Planet 214
- 10 Mars: A Near Miss for Life? 233
- 11 Jupiter: Giant of the Solar System 260
- 12 Saturn: Spectacular Rings and Mysterious Moons 286
- 13 Uranus and Neptune: The Outer Worlds of the Solar System 314
- 14 Solar System Debris: Keys to Our Origin 333
- 15 Exoplanets: Planetary Systems Beyond Our Own 363

PART III: Stars and Stellar Evolution 384

- 16 The Sun: Our Parent Star 386
- 17 The Stars: Giants, Dwarfs, and the Main Sequence 417
- 18 The Interstellar Medium: Gas and Dust Among the Stars 445
- 19 Star Formation: A Traumatic Birth 465
- 20 Stellar Evolution: The Life and Death of a Star 490
- 21 Stellar Explosions: Novae, Supernovae, and the Formation of the Elements 516
- 22 Neutron Stars and Black Holes: Strange States of Matter 537

PART IV: Galaxies and Cosmology 570

- 23 The Milky Way Galaxy: A Spiral in Space 572
- 24 Galaxies: Building Blocks of the Universe 601
- 25 Galaxies and Dark Matter: The Large-Scale Structure of the Cosmos 632
- 26 Cosmology: The Big Bang and the Fate of the Universe 659
- 27 The Early Universe: Toward the Beginning of Time 681
- 28 Life in the Universe: Are We Alone? 706

CONTENTS

About the Authors xvi

Preface xvii

PART I: Astronomy and the Universe 2

1 Charting the Heavens

The Foundations of Astronomy 5

- 1.1 Our Place in Space 6
 - 1.2 Scientific Theory and the Scientific Method 8
 - 1.3 The “Obvious” View 10
 - 1.4 Earth’s Orbital Motion 13
 - MORE PRECISELY 1-1 Angular Measure 14
 - 1.5 The Motion of the Moon 18
 - 1.6 The Measurement of Distance 24
 - MORE PRECISELY 1-2 Measuring Distances with Geometry 28
 - Chapter Review 29
-

2 The Copernican Revolution

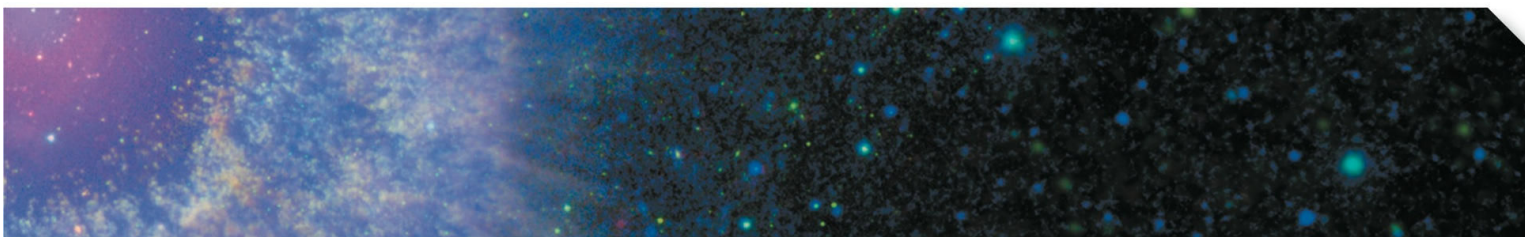
The Birth of Modern Science 33

- 2.1 Ancient Astronomy 34
 - 2.2 The Geocentric Universe 36
 - 2.3 The Heliocentric Model of the Solar System 40
 - DISCOVERY 2-1 Foundations of the Copernican Revolution 40
 - 2.4 The Birth of Modern Astronomy 41
 - 2.5 The Laws of Planetary Motion 44
 - MORE PRECISELY 2-1 Some Properties of Planetary Orbits 46
 - 2.6 The Dimensions of the Solar System 48
 - 2.7 Newton’s Laws 49
 - 2.8 Newtonian Mechanics 52
 - MORE PRECISELY 2-2 Weighing the Sun 53
 - Chapter Review 56
-

3 Radiation

Information from the Cosmos 59

- 3.1 Information from the Skies 60
- 3.2 Waves in What? 63
- 3.3 Electromagnetic Spectrum 65
- 3.4 Thermal Radiation 67
 - DISCOVERY 3-1 The Wave Nature of Radiation 68
 - MORE PRECISELY 3-1 The Kelvin Temperature Scale 69
 - MORE PRECISELY 3-2 More About the Radiation Laws 72
- 3.5 The Doppler Effect 73
 - MORE PRECISELY 3-3 Measuring Velocities with the Doppler Effect 75
- Chapter Review 76



4 Spectroscopy

The Inner Workings of Atoms 79

- 4.1 Spectral Lines 80
- 4.2 Atoms and Radiation 84
 - DISCOVERY 4-1 The Photoelectric Effect 85
 - MORE PRECISELY 4-1 The Hydrogen Atom 87
- 4.3 Formation of Spectral Lines 88
- 4.4 Molecules 91
- 4.5 Spectral-Line Analysis 92
- Chapter Review 96

5 Telescopes

The Tools of Astronomy 99

- 5.1 Optical Telescopes 100
- 5.2 Telescope Size 105
- 5.3 Images and Detectors 109
- 5.4 High-Resolution Astronomy 111
- 5.5 Radio Astronomy 114
- 5.6 Interferometry 118
- 5.7 Space-Based Astronomy 121
 - DISCOVERY 5-1 The ALMA Array 122
- 5.8 Full-Spectrum Coverage 128
- Chapter Review 129

PART II: Our Planetary System 132

6 The Solar System

Comparative Planetology and Formation Models 134

- 6.1 An Inventory of the Solar System 135
- 6.2 Measuring the Planets 137
- 6.3 The Overall Layout of the Solar System 139
- 6.4 Terrestrial and Jovian Planets 140
- 6.5 Interplanetary Matter 141
 - DISCOVERY 6-1 Gravitational "Slingshots" 142
- 6.6 How Did the Solar System Form? 144
 - DISCOVERY 6-2 Spacecraft Exploration of the Solar System 146
 - MORE PRECISELY 6-1 Angular Momentum 149
- 6.7 Jovian Planets and Planetary Debris 152
- Chapter Review 156

7 Earth

Our Home in Space 160

- 7.1 Overall Structure of Planet Earth 161

7.2	Earth's Atmosphere	162
	MORE PRECISELY 7-1 Why Is the Sky Blue?	164
	DISCOVERY 7-1 The Greenhouse Effect and Global Warming	166
7.3	Earth's Interior	167
	MORE PRECISELY 7-2 Radioactive Dating	170
7.4	Surface Activity	172
7.5	Earth's Magnetosphere	178
7.6	The Tides	180
	Chapter Review	183

8 The Moon and Mercury

Scorched and Battered Worlds 187

8.1	Orbital Properties	188
8.2	Physical Properties	189
8.3	Surface Features	190
8.4	Rotation Rates	193
	MORE PRECISELY 8-1 Why Air Sticks Around	194
	DISCOVERY 8-1 Lunar Exploration	196
8.5	Lunar Cratering and Surface Composition	199
8.6	The Surface of Mercury	204
8.7	Interiors	206
8.8	Origin of the Moon	207
8.9	Evolutionary History of the Moon and Mercury	209
	Chapter Review	211

9 Venus

Earth's Sister Planet 214

9.1	Orbital Properties	215
9.2	Physical Properties	216
9.3	Long-Distance Observations of Venus	217
9.4	The Surface of Venus	218
9.5	The Atmosphere of Venus	225
9.6	Venus's Magnetic Field and Internal Structure	229
	Chapter Review	230

10 Mars

A Near Miss for Life? 233

10.1	Orbital Properties	234
10.2	Physical Properties	235
10.3	Long-Distance Observations of Mars	236
10.4	The Martian Surface	236
10.5	Water on Mars	240
	DISCOVERY 10-1 Life on Mars?	246
10.6	The Martian Atmosphere	252

- 10.7 Martian Internal Structure 255
 - 10.8 The Moons of Mars 255
 - Chapter Review 257
-

11 Jupiter

- Giant of the Solar System 260
 - 11.1 Orbital and Physical Properties 261
 - 11.2 Jupiter's Atmosphere 263
 - 11.3 Internal Structure 269
 - DISCOVERY 11-1 A Cometary Impact 270
 - 11.4 Jupiter's Magnetosphere 271
 - DISCOVERY 11-2 Almost a Star? 274
 - 11.5 The Moons of Jupiter 274
 - 11.6 Jupiter's Ring 282
 - Chapter Review 283
-

12 Saturn

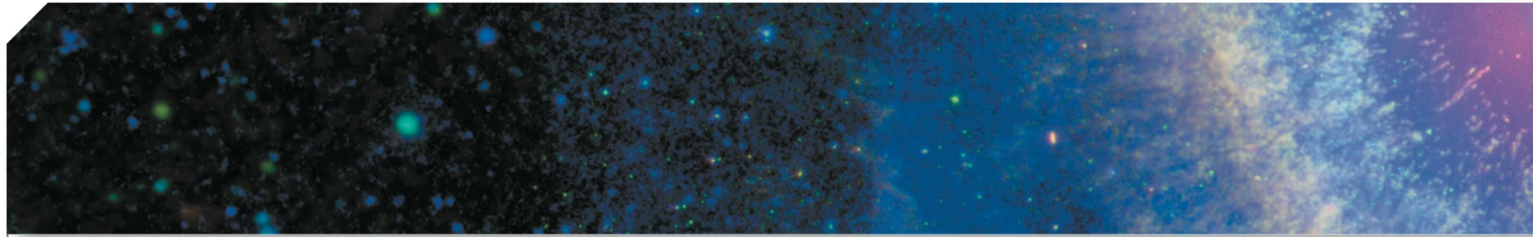
- Spectacular Rings and Mysterious Moons 286
 - 12.1 Orbital and Physical Properties 287
 - 12.2 Saturn's Atmosphere 288
 - 12.3 Saturn's Interior and Magnetosphere 292
 - 12.4 Saturn's Spectacular Ring System 293
 - 12.5 The Moons of Saturn 300
 - DISCOVERY 12-1 Dancing Among Saturn's Moons 302
 - Chapter Review 310
-

13 Uranus and Neptune

- The Outer Worlds of the Solar System 314
 - 13.1 The Discoveries of Uranus and Neptune 315
 - 13.2 Orbital and Physical Properties 317
 - 13.3 The Atmospheres of Uranus and Neptune 319
 - 13.4 Magnetospheres and Internal Structure 321
 - 13.5 The Moon Systems of Uranus and Neptune 323
 - 13.6 The Rings of the Outermost Jovian Planets 326
 - Chapter Review 330
-

14 Solar System Debris

- Keys to Our Origin 333
- 14.1 Asteroids 334
- 14.2 Comets 340
 - DISCOVERY 14-1 What Killed the Dinosaurs? 346
- 14.3 Beyond Neptune 348
- 14.4 Meteoroids 355
- Chapter Review 359



15 Exoplanets

Planetary Systems Beyond Our Own 363

- 15.1 Modeling Planet Formation 364
- 15.2 Solar System Regularities and Irregularities 365
- 15.3 Exoplanet Detection 366
- 15.4 Exoplanet Properties 370
 - DISCOVERY 15-1 Gravitational Microlensing 372
- 15.5 Is Our Solar System Unusual? 376
 - Chapter Review 380

PART III: Stars and Stellar Evolution 384

16 The Sun

Our Parent Star 386

- 16.1 Physical Properties of the Sun 387
- 16.2 The Solar Interior 389
 - DISCOVERY 16-1 Eavesdropping on the Sun 392
- 16.3 The Sun's Atmosphere 394
- 16.4 Solar Magnetism 397
- 16.5 The Active Sun 402
 - DISCOVERY 16-2 Solar-Terrestrial Relations 406
- 16.6 The Heart of the Sun 407
 - MORE PRECISELY 16-1 Fundamental Forces 410
- 16.7 Observations of Solar Neutrinos 410
 - MORE PRECISELY 16-2 Energy Generation in the Proton-Proton Chain 412
 - Chapter Review 413

17 The Stars

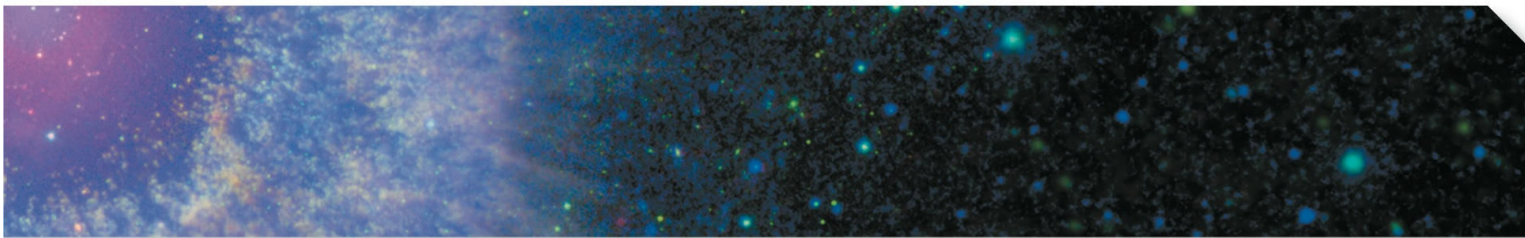
Giants, Dwarfs, and the Main Sequence 417

- 17.1 The Solar Neighborhood 418
- 17.2 Luminosity and Apparent Brightness 421
- 17.3 Stellar Temperatures 424
 - MORE PRECISELY 17-1 More on the Magnitude Scale 427
- 17.4 Stellar Sizes 428
- 17.5 The Hertzsprung-Russell Diagram 430
 - MORE PRECISELY 17-2 Estimating Stellar Radii 431
- 17.6 Extending the Cosmic Distance Scale 433
- 17.7 Stellar Masses 436
- 17.8 Mass and Other Stellar Properties 438
 - MORE PRECISELY 17-3 Measuring Stellar Masses in Binary Stars 439
 - Chapter Review 441

18 The Interstellar Medium

Gas and Dust Among the Stars 445

- 18.1 Interstellar Matter 446

- 
- 18.2 Emission Nebulae 449
 - 18.3 Dark Dust Clouds 455
 - 18.4 21-Centimeter Radiation 458
 - 18.5 Interstellar Molecules 459
 - Chapter Review 462
-

19 Star Formation

- A Traumatic Birth 465
 - 19.1 Star-Forming Regions 466
 - MORE PRECISELY 19-1 Competition in Star Formation 468
 - 19.2 Formation of Stars Like the Sun 468
 - 19.3 Stars of Other Masses 473
 - 19.4 Observations of Cloud Fragments and Protostars 474
 - DISCOVERY 19-1 Observations of Brown Dwarfs 475
 - 19.5 Shock Waves and Star Formation 480
 - 19.6 Star Clusters 481
 - DISCOVERY 19-2 Eta Carinae 486
 - Chapter Review 487
-

20 Stellar Evolution

- The Life and Death of a Star 490
 - 20.1 Leaving the Main Sequence 491
 - 20.2 Evolution of a Sun-Like Star 492
 - 20.3 The Death of a Low-Mass Star 497
 - DISCOVERY 20-1 Learning Astronomy from History 503
 - 20.4 Evolution of Stars More Massive than the Sun 504
 - DISCOVERY 20-2 Mass Loss from Giant Stars 506
 - 20.5 Observing Stellar Evolution in Star Clusters 507
 - 20.6 Stellar Evolution in Binary Systems 510
 - Chapter Review 513
-

21 Stellar Explosions

- Novae, Supernovae, and the Formation of the Elements 516
 - 21.1 Life After Death for White Dwarfs 517
 - 21.2 The End of a High-Mass Star 519
 - 21.3 Supernovae 521
 - 21.4 Formation of the Elements 525
 - DISCOVERY 21-1 Supernova 1987A 526
 - 21.5 The Cycle of Stellar Evolution 533
 - Chapter Review 534
-

22 Neutron Stars and Black Holes

- Strange States of Matter 537
- 22.1 Neutron Stars 538
- 22.2 Pulsars 539

- 22.3 Neutron-Star Binaries 542
- 22.4 Gamma-Ray Bursts 546
- 22.5 Black Holes 549
- 22.6 Einstein's Theories of Relativity 552
 - DISCOVERY 22-1 Special Relativity 553
- 22.7 Space Travel Near Black Holes 556
- 22.8 Observational Evidence for Black Holes 559
 - MORE PRECISELY 22-1 Tests of General Relativity 560
 - DISCOVERY 22-2 Gravitational Waves: A New Window on the Universe 564
- Chapter Review 567

PART IV: Galaxies and Cosmology 570

23 The Milky Way Galaxy

A Spiral in Space 572

- 23.1 Our Parent Galaxy 573
- 23.2 Measuring the Milky Way 575
 - DISCOVERY 23-1 Early "Computers" 579
- 23.3 Galactic Structure 581
- 23.4 Formation of the Milky Way 584
- 23.5 Galactic Spiral Arms 586
 - DISCOVERY 23-2 Density Waves 589
- 23.6 The Mass of the Milky Way Galaxy 590
- 23.7 The Galactic Center 593
- Chapter Review 598

24 Galaxies

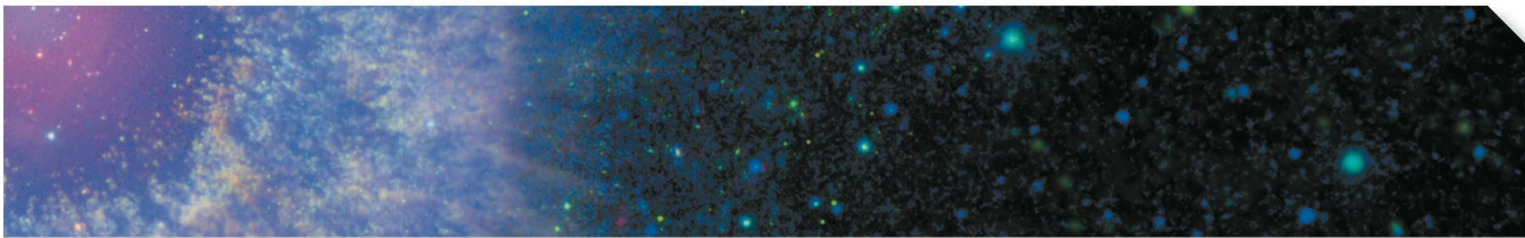
Building Blocks of the Universe 601

- 24.1 Hubble's Galaxy Classification 602
- 24.2 The Distribution of Galaxies in Space 608
- 24.3 Hubble's Law 612
- 24.4 Active Galactic Nuclei 615
 - MORE PRECISELY 24-1 Relativistic Redshifts and Look-Back Time 616
- 24.5 The Central Engine of an Active Galaxy 624
- Chapter Review 628

25 Galaxies and Dark Matter

The Large-Scale Structure of the Cosmos 632

- 25.1 Dark Matter in the Universe 633
- 25.2 Galaxy Collisions 636
- 25.3 Galaxy Formation and Evolution 638
 - DISCOVERY 25-1 The Sloan Digital Sky Survey 644
- 25.4 Black Holes in Galaxies 645
- 25.5 The Universe on Large Scales 649
- Chapter Review 656



26 Cosmology

The Big Bang and the Fate of the Universe 659

- 26.1 The Universe on the Largest Scales 660
- 26.2 The Expanding Universe 662
- 26.3 The Fate of the Cosmos 665
- 26.4 The Geometry of Space 668
 - MORE PRECISELY 26-1 Curved Space 669
- 26.5 Will the Universe Expand Forever? 670
 - DISCOVERY 26-1 Einstein and the Cosmological Constant 672
- 26.6 Dark Energy and Cosmology 673
- 26.7 The Cosmic Microwave Background 675
- Chapter Review 678

27 The Early Universe

Toward the Beginning of Time 681

- 27.1 Back to the Big Bang 682
- 27.2 Evolution of the Universe 685
 - MORE PRECISELY 27-1 More on Fundamental Forces 687
- 27.3 Formation of Nuclei and Atoms 689
- 27.4 The Inflationary Universe 692
- 27.5 Large-Scale Structure in the Universe 696
- 27.6 Cosmic Structure and the Microwave Background 699
- Chapter Review 703

28 Life in the Universe

Are We Alone? 706

- 28.1 Cosmic Evolution 707
 - DISCOVERY 28-1 The Virus 709
- 28.2 Life in the Solar System 713
- 28.3 Intelligent Life in the Galaxy 715
- 28.4 The Search for Extraterrestrial Intelligence 720
- Chapter Review 724

Appendices

- APPENDIX 1 Scientific Notation A-1
- APPENDIX 2 Astronomical Measurement A-2
- APPENDIX 3 Tables A-3
- GLOSSARY G-1
- ANSWERS TO CHECK QUESTIONS AK-1
- ANSWERS TO SELF-TEST QUESTIONS AK-7
- PHOTO CREDITS/TEXT PERMISSIONS C-1
- INDEX I-1
- STAR CHARTS S-1

ONLINE CONTENTS

PART I: Astronomy and the Universe **2**

Chapter 1 Charting the Heavens 4

- SELF-GUIDED TUTORIAL Phases of the Moon 19
- NARRATED FIGURE Lunar Phases 19

Chapter 2 The Copernican Revolution 33

- INTERACTIVE FIGURE Geocentric Model 38
- INTERACTIVE FIGURE Retrograde Motion 41
- INTERACTIVE FIGURE Venus Phases 43

Chapter 3 Radiation 59

- NARRATED INTERACTIVE FIGURE Electromagnetic Spectrum 67
- SELF-GUIDED TUTORIAL Continuous Spectra and Blackbody Radiation 70
- SELF-GUIDED TUTORIAL Doppler Effect 74

Chapter 4 Spectroscopy 79

- SELF-GUIDED TUTORIAL Emission Spectra 82
- SELF-GUIDED TUTORIAL Absorption Spectra 82

Chapter 5 Telescopes 99

- SELF-GUIDED TUTORIAL The Optics of a Simple Lens 101
- SELF-GUIDED TUTORIAL Chromatic Aberration 102
- SELF-GUIDED TUTORIAL Reflecting Telescopes 103
- NARRATED FIGURE Multiple Wavelengths 128

PART II: Our Planetary System **132**

Chapter 6 The Solar System 134

- INTERACTIVE FIGURE Nebular Contraction 145

Chapter 7 Earth 160

- SELF-GUIDED TUTORIAL The Greenhouse Effect 166
- INTERACTIVE FIGURE Solar and Lunar Tides 182

Chapter 8 The Moon and Mercury 187

- NARRATED FIGURE Moon Formation 208

Chapter 9 Venus 214

- NARRATED FIGURE Venus's Brightness 215

- SELF-GUIDED TUTORIAL Super-spaceship—Voyage to Venus 219

Chapter 10 Mars 233

- INTERACTIVE FIGURE Mars Map 237
- SELF-GUIDED TUTORIAL Comparative Planetology: Mars 245
- SELF-GUIDED TUTORIAL Atmospheric Lifetimes 255

Chapter 11 Jupiter 260

- SELF-GUIDED TUTORIAL Jupiter—Differential Rotation 263
- INTERACTIVE FIGURE Galilean Moons 275

Chapter 12 Saturn 286

- NARRATED FIGURE Roche Limit 295
- INTERACTIVE FIGURE Saturn's Rings, Up Close 297

Chapter 13 Uranus and Neptune 314

- NARRATED FIGURE Jovian Magnetic Fields 322

Chapter 14 Solar System Debris 333

- NARRATED FIGURE Inner Solar System 334



PART III: Stars and Stellar Evolution **384**

Chapter 16 The Sun 386

- SELF-GUIDED TUTORIAL Super-Spaceship—Voyage to the Sun **388**
- NARRATED FIGURE Stellar Balance **390**

Chapter 17 The Stars 417

- SELF-GUIDED TUTORIAL Stellar Parallax **418**
- NARRATED FIGURE Inverse-Square Law **422**
- SELF-GUIDED TUTORIAL Hertzsprung-Russell Diagram **430**
- SELF-GUIDED TUTORIAL Binary Stars—Radial Velocity Curves **436**
- SELF-GUIDED TUTORIAL Eclipsing Binary Stars—Light Curves **437**

Chapter 18 The Interstellar Medium 445

- NARRATED FIGURE Reddening **447**

Chapter 19 Star Formation 465

- INTERACTIVE FIGURE Newborn Star on the H-R Diagram **472**

Chapter 20 Stellar Evolution 490

- SELF-GUIDED TUTORIAL Evolution of a 1-Solar-Mass Star **504**

Chapter 22 Neutron Stars and Black Holes 537

- NARRATED FIGURE Pulsar Model **540**
- SELF-GUIDED TUTORIAL Escape Speed and Black Hole Event Horizons **551**
- INTERACTIVE FIGURE Gravitational Redshift **557**

PART IV: Galaxies and Cosmology **570**

Chapter 23 The Milky Way Galaxy 572

- NARRATED FIGURE Globular Cluster Distribution **580**

- SELF-GUIDED TUTORIAL Gravitational Lensing **592**
- INTERACTIVE FIGURE Galactic Center **593**

Chapter 24 Galaxies 601

- NARRATED FIGURE Galaxy Rotation **610**
- INTERACTIVE FIGURE M87 Jet **621**

Chapter 25 Galaxies and Dark Matter 632

- INTERACTIVE FIGURE Starburst Galaxy **637**

Chapter 26 Cosmology 659

- INTERACTIVE FIGURE The Expanding Raisin Cake (Universe) **664**

Chapter 27 The Early Universe 681

- NARRATED FIGURE Structure Formation **698**
- INTERACTIVE FIGURE Early Structure **700**

Chapter 28 Life in the Universe 706

- NARRATED FIGURE Drake Equation **716**

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PREFACE

Astronomy is a science that thrives on new discoveries. Fueled by new technologies and novel theoretical insights, the study of the cosmos continues to change our understanding of the universe. We are pleased to have the opportunity to present in this book a representative sample of the known facts, evolving ideas, and frontier discoveries in astronomy today.

Astronomy Today has been written for students who have taken no previous college science courses and who will likely not major in physics or astronomy. It is intended for use in a one- or two-semester, nontechnical astronomy course. We present a broad view of astronomy, straightforwardly descriptive and without complex mathematics. The absence of sophisticated mathematics, however, in no way prevents discussion of important concepts. Rather, we rely on qualitative reasoning as well as analogies with objects and phenomena familiar to the student to explain the complexities of the subject without oversimplification. We have tried to communicate the excitement we feel about astronomy and to awaken students to the marvelous universe around us.

We are very gratified that the first eight editions of this text have been so well received by many in the astronomy education community. In using those earlier texts, many teachers and students have given us helpful feedback and constructive criticisms. From these, we have learned to communicate better both the fundamentals and the excitement of astronomy. Many improvements inspired by these comments have been incorporated into this new edition.

Focus of the Ninth Edition

From the first edition, we have tried to meet the challenge of writing a book that is both accurate and approachable. To the student, astronomy sometimes seems like a long list of unfamiliar terms to be memorized and repeated. Many new terms and concepts are introduced in this course, but we hope students will also learn and remember how science is done, how the universe works, and how things are connected. In the ninth edition, we have taken particular care to show how astronomers know what they know, and to highlight both the scientific principles underlying their work and the process used in discovery.

New and Revised Material

Astronomy is a rapidly evolving field and, in the three years since the publication of the eighth edition of *Astronomy Today*, has seen many new discoveries covering the entire spectrum of astronomical research. Almost every chapter in the ninth edition has been substantially updated with new information.

Several chapters have also seen significant reorganization in order to streamline the overall presentation, strengthen our focus on the process of science, and reflect new understanding and emphases in contemporary astronomy.

In addition to updates throughout the text on the numbers and properties of the many astronomical objects, the many substantive changes include:

- New discussion in Chapter 5 of next-generation telescopes and high-resolution astronomy.
- Updated information and imagery in *Discovery 5-1* on the ALMA array.
- New discussion in Chapter 8 of ice on the Moon.
- Additional coverage in Chapter 8 of Mercury's surface and interior based on *Messenger* data.
- New discussion in Chapter 10 of the depletion of the Martian atmosphere.
- Expanded coverage in Chapter 10 of the *Curiosity* rover on Mars and its findings so far.
- Update in Chapter 11 on the changing appearance of Jupiter's Great Red Spot.
- New material in Chapter 11 on the 2016 *Juno* mission.
- Updated discussion in Chapter 11 of the internal structure of Ganymede.
- New discussion in Chapter 12 of storms on Saturn.
- Expanded coverage in Chapter 12 of lakes and other features on the surface of Saturn's moon Titan.
- Additional material in Chapter 12 on Saturn's moon Enceladus.
- Update in Chapter 13 on the return of Neptune's Dark Spot.
- New coverage in Chapter 14 of the *Dawn* mission to Ceres.
- Extensive discussion in Chapter 14 of the *Rosetta* mission to comet 67 P/Churyumov-Gerasimenko.
- Completely revised presentation of Pluto in Chapter 14 based on data from the *New Horizons* mission.
- Updated and rewritten presentation in Chapter 15 of exoplanet searches and properties.
- New material in Chapter 15 on direct imaging as an exoplanet detection technique.
- Expanded discussion in Chapter 15 of exoplanet composition.
- Presentation in new *Discovery 15-1* of gravitational microlensing as an important exoplanet detection technique.
- Updated discussion in Chapter 15 of habitable zones and planetary systems in star clusters.

- Additional material in *Discovery 16-1 on the Solar Dynamics Observatory*.
- Substantially improved discussion and imagery in Chapter 16 of the sunspot cycle.
- New coverage in Chapter 17 on the *GAIA* astrometric mission.
- Additional material in *Discovery 19-1* on brown dwarfs, clarifying the distinction between brown dwarfs and planets.
- Improved text and imagery in Chapter 19 on the observational evidence for various stages of star formation: ALMA imagery of protostellar collapse; *HST* observations of protoplanetary disks.
- Improved simulations of star cluster formation in Chapter 19.
- Expanded discussion in *Discovery 20-2* of mass loss from giant stars.
- Emphasis in Chapter 22 of the connection between hypernovae and black holes.
- Expanded treatment in *Discovery 22-1* of relativity and time dilation.
- Rewritten discussion of gravitational radiation in *Discovery 22-1*, including extensive coverage of the 2015 LIGO detections.
- Reconsideration in Chapter 22 of the existence of black holes, in the light of the new LIGO findings.
- New coverage in Chapter 23 of the “X” in the Milky Way bulge and its implication for our Galaxy’s history.
- Expanded material in Chapter 23 on the “S stars” in the Galactic center and energetic outflows from the Galactic center into the halo.
- Updated discussion of extremophilic life in Chapter 28.
- Consistent distance scales in all figures, helping students gain an understanding of the vastness of the universe.
- Numerous replacement images for currency and clarity, and updated art throughout the text.

The Illustration Program

Visualization plays an important role in both the teaching and the practice of astronomy, and we continue to place strong emphasis on this aspect of our book. We have tried to combine aesthetic beauty with scientific accuracy in the artist’s conceptions that adorn the text, and we have sought to present the best and latest imagery of a wide range of cosmic objects. Each illustration has been carefully crafted to enhance student learning; each is pedagogically sound and tied tightly to the nearby discussion of important scientific facts and ideas. This edition contains more than 100 revised figures that show the latest imagery and the results learned from them.

Compound Art It is rare that a single image, be it a photograph or an artist’s conception, can capture all aspects of a complex subject. Wherever possible, multiple-part figures are used in an attempt to convey the greatest amount of information in the most vivid way:

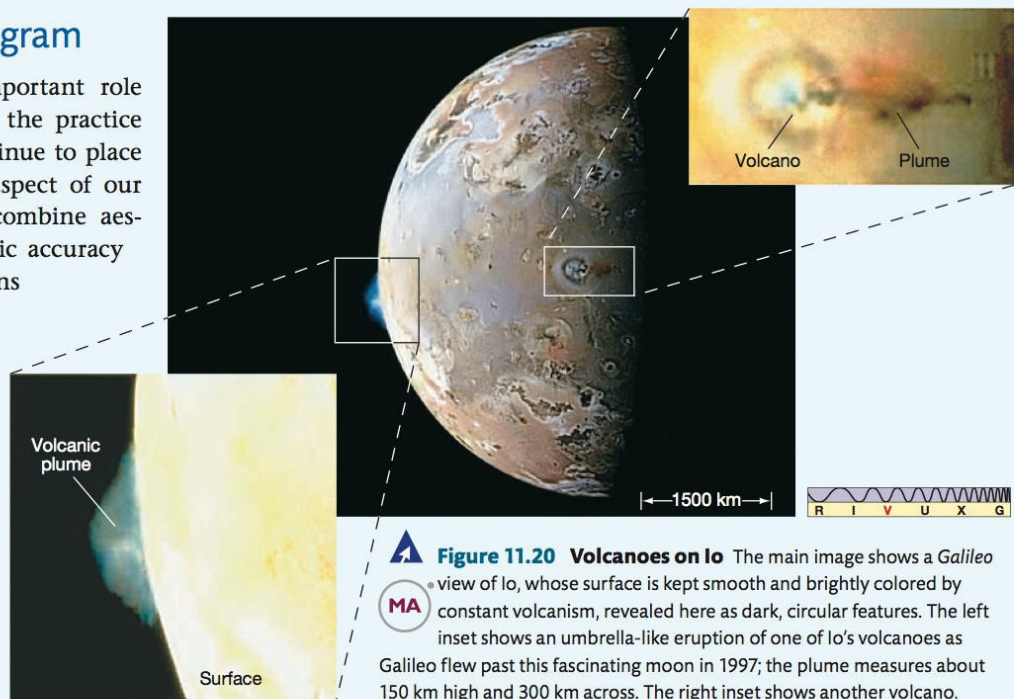


Figure 11.20 Volcanoes on Io The main image shows a *Galileo* view of Io, whose surface is kept smooth and brightly colored by constant volcanism, revealed here as dark, circular features. The left inset shows an umbrella-like eruption of one of Io’s volcanoes as *Galileo* flew past this fascinating moon in 1997; the plume measures about 150 km high and 300 km across. The right inset shows another volcano, this one face-on, where surface features here are resolved to just a few kilometers. (NASA)

- Visible images are often presented along with their counterparts captured at other wavelengths.
- Interpretive line drawings are often superimposed on or juxtaposed with real astronomical photographs, helping students to really “see” what the photographs reveal.

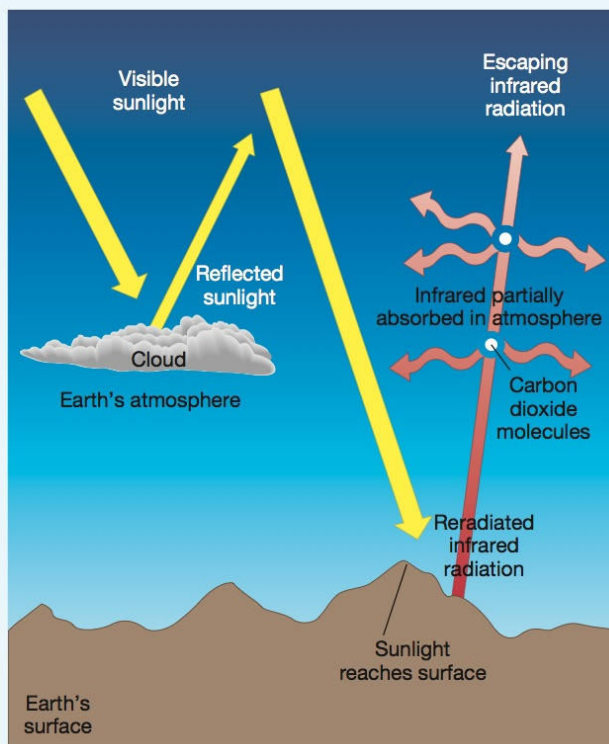


Figure 7.5 Greenhouse Effect Sunlight that is not reflected by clouds reaches Earth's surface, warming it up. Infrared radiation reradiated from the surface is partially absorbed by carbon dioxide (and also water vapor, not shown here) in the atmosphere, causing the overall surface temperature to rise.

- Breakouts—often multiple ones—are used to zoom in from wide-field shots to close-ups so that detailed images can be understood in their larger context.

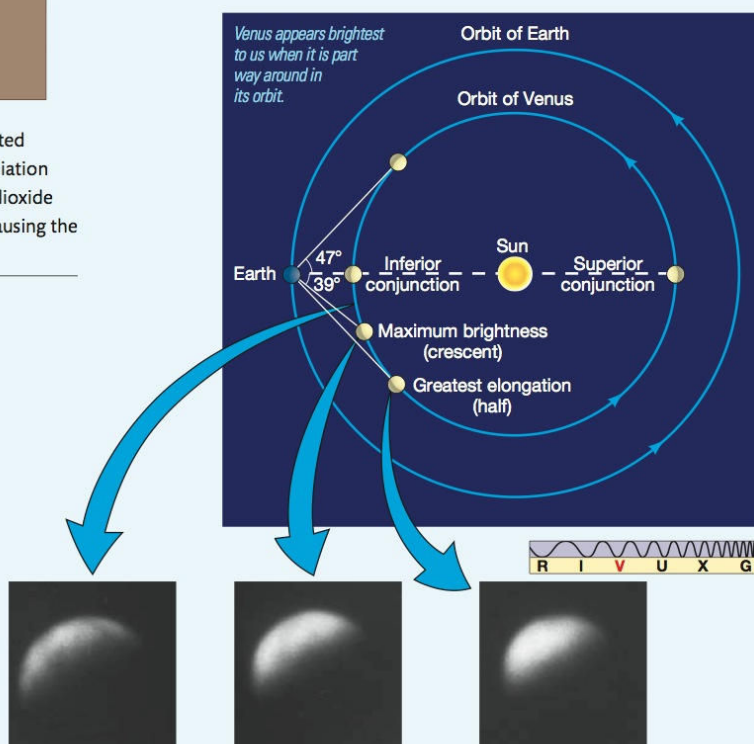
Figures and Photos Icons throughout the text direct students to dynamic, interactive versions of art and photos on MasteringAstronomy®. Using online applets, students can manipulate factors such as time, wavelength, scale, and perspective to increase their understanding of these figures.

NARRATED Figures Narrated Figures are brief videos that step students through complex figures from the text, expanding students' understanding of fundamental concepts in a presentation that includes narration, enhanced visuals, and one to two embedded questions, followed by short, one- to two-question Mastering activities that are graded. They mirror how an instructor might present a topic in class and can be assigned as homework, self-study, or as part of a pre-lecture program.

Figure Annotations The ninth edition incorporates the research-proven technique of strategically placing annotations (which always appear in blue type) within key pieces of art, fostering students' ability to read and interpret complex figures, focus on the most relevant information, and integrate written and visual knowledge.

Full Spectrum Coverage and Spectrum Icons

Astronomers exploit the full range of the electromagnetic spectrum to gather information about the cosmos. Throughout this book, images taken at radio, infrared, ultraviolet, X-ray, or gamma-ray wavelengths are used to supplement visible-light images. As it is sometimes difficult (even for a professional) to tell at a glance which images are visible-light photographs and which are false-color images created with other wavelengths, each photo in the text is accompanied by an icon that identifies the wavelength of electromagnetic radiation used to capture the image.



NARRATED Figure 9.2 Venus's Brightness Venus appears full when it is at its greatest distance from Earth, on the opposite side of the Sun from us (superior conjunction). As its distance decreases, less and less of its sunlit side becomes visible. When closest to Earth, it lies between us and the Sun (inferior conjunction), so we cannot see the sunlit side of the planet at all. Venus appears brightest when it is about 39° from the Sun. (Compare Figure 2.12.) (Insets: UC Regents/Lick Observatory)

Other Pedagogical Features

As with many other parts of our text, instructors have helped guide us toward what is most helpful for effective student learning. With their assistance, we have revised both our in-chapter and end-of-chapter pedagogical apparatus to increase its utility to students.

Learning Outcomes Studies indicate that beginning students have trouble prioritizing textual material. For this reason, a few (typically five or six) well-defined Learning

Outcomes are provided at the start of each chapter. These help students structure their reading of the chapter and then test their mastery of key concepts. The Learning Outcomes are numbered and keyed to the items in the Chapter Summary, which in turn refer back to passages in the text. This highlighting of the most important aspects of the chapter helps students prioritize information and also aids in their review. The Learning Outcomes are organized and phrased in such a way as to make them objectively testable, affording students a means of gauging their own progress.

LEARNING OUTCOMES

Studying this chapter will enable you to

- 1 Summarize the composition and physical properties of the interstellar medium.
- 2 Describe the characteristics of emission nebulae, and explain their significance in the life cycle of stars.
- 3 List the basic properties of dark interstellar clouds.
- 4 Specify the radio techniques used to probe the nature of interstellar matter.
- 5 Explain the nature and significance of interstellar molecules.

The Big Picture The Big Picture feature on every chapter opening spread encapsulates the overarching message that each chapter imparts, helping students see how chapter

content is connected to a broad understanding of the universe.

The Big Picture

Stars are everywhere in the night sky. The naked eye can distinguish about 6000 of them, spread across 88 constellations. With binoculars or even a small telescope, millions more become visible. The total number of stars, even in our local cosmic neighborhood, is virtually beyond our ability to count. By analyzing the light from millions of distant stars, astronomers have learned a great deal about stellar properties—their masses and radii, their luminosities, even their ages and destinies. Stars tell us more about the fundamentals of astronomy than any other class of objects in the universe.

The Big Question Each chapter now ends with a broad, open-ended query that is intended to ignite students' curiosity about the still-unanswered questions at the forefront

of astronomical research. The Big Question builds on the material presented in the chapter and invites students to speculate on the larger scope of what they have just learned.

The Big Question

We will see in the next few chapters that, once you specify the mass and composition of a star, its structure and future evolution are largely set. But are the masses of newly formed stars pretty much the same everywhere, or do they vary systematically from place to place in our Galaxy and beyond? Astronomers generally assume the former, since that's the only practical way we can do astronomy, but is it true? At the cutting edge of research, astronomers are now testing this fundamental assumption against state-of-the-art observations.

Concept Checks We incorporate into each chapter a number of "Concept Checks"—key questions that require the reader to reconsider some of the material just presented or attempt to place it into a broader context. Answers to these in-chapter questions are provided at the back of the book.

CONCEPT Check

- ▶ Why do astronomers draw such a clear distinction between the inner and the outer planets?


Process of Science Checks Each chapter now also includes one or two “Process of Science Checks,” similar to the Concept Checks but aimed specifically at clarifying the questions of how science is done and how scientists reach the conclusions they do. Answers to these in-chapter questions are also provided at the back of the book.

PROCESS OF SCIENCE Check

► In what sense are the comets we see *unrepresentative* of comets in general?

Data Points (NEW) Data Points sidebars in each chapter, based on data captured from thousands of students, alert students to the statistically most common mistakes made when working problems on related topics in MasteringAstronomy*.

Concept Links In astronomy, as in many scientific disciplines, almost every topic seems to have some bearing on almost every other. In particular, the connection between the astronomical material and the physical principles set forth early in the text is crucial. Practically everything in Chapters 6–28 of this text rests on the foundation laid in the first five chapters. For example, it is important that students, when they encounter the discussion of high-redshift objects in Chapter 25, recall not only what they just learned about Hubble’s law in Chapter 24 but also refresh their memories, if necessary, about the inverse-square law (Chapter 17), stellar spectra (Chapter 4), and the Doppler shift (Chapter 3). Similarly, the discussions of the mass of binary-star components (Chapter 17) and of galactic rotation (Chapter 23) both depend on the discussion of Kepler’s and Newton’s laws in Chapter 2. Throughout, discussions of new astronomical objects and concepts rely heavily on comparison with topics introduced earlier in the text.

It is important to remind students of these links so that they recall the principles on which later discussions rest and, if necessary, review them. To this end, we have inserted “concept links” throughout the text—symbols that mark key intellectual bridges between material in different chapters. The links, denoted by the symbol  together with a section reference, signal that the topic under discussion is related in some significant way to ideas developed earlier and provide direction to material to review before proceeding.

Key Terms Like all subjects, astronomy has its own specialized vocabulary. To aid student learning, the most important astronomical terms are boldfaced at their first appearance in the text. Boldfaced Key Terms in the Chapter Summary are linked with the page number where the term was defined. In addition, an expanded alphabetical glossary, defining each Key Term and locating its first use in the text, appears at the end of the book.

H–R Diagrams and Acetate Overlays All of the book’s H–R diagrams are drawn in a uniform format, using real data. In addition, a unique set of transparent acetate overlays dramatically demonstrates to students how the H–R diagram helps us to organize our information about the stars and track their evolutionary histories.

More Precisely Boxes These boxes provide more quantitative treatments of subjects discussed qualitatively in the text. Removing these more challenging topics from the main flow of the narrative and placing them within a separate modular element of the chapter design (so that they can be covered in class, assigned as supplementary material, or simply left as optional reading for those students who find them of interest) will allow instructors greater flexibility in setting the level of their coverage.

Discovery Boxes Exploring a wide variety of interesting supplementary topics, Discovery boxes provide the reader with insight into how scientific knowledge evolves and emphasizes the process of science.

End-of-Chapter Questions, Problems, and Activities (Extensively Revised)

- Each chapter incorporates **Review and Discussion Questions**, which may be used for in-class review or for assignment. As with the Self-Test Questions, the material needed to answer Review Questions may be found within the chapter. The Discussion Questions explore particular topics more deeply, often asking for opinions, not just facts. As with all discussions, these questions usually have no single “correct” answer. Questions identified with a **POS** icon encourage students to explore the Process of Science, and each Learning Outcome is reflected in one of the Review and Discussion questions, marked by **LO**. **BP (NEW)** questions relate to the Big Picture item at the start of the chapter.
- Each chapter also contains **Conceptual Self-Test Questions** in a multiple-choice format, including select questions that are tied directly to a specific figure or diagram in the text, allowing students to assess their understanding of the chapter material. These questions are identified with a **vis** icon. Answers to all these questions appear at the end of the book.
- The end-of-chapter material includes **Problems**, based on the chapter contents and requiring some numerical calculation. In many cases the problems are tied directly to quantitative statements made (but not worked out in detail) in the text. The solutions to the problems are not contained verbatim within the chapter, but the information necessary to solve them has been presented in the text. Answers to odd-numbered Problems appear at the end of the book.
- Heavily revised in this edition, the end-of-chapter material now ends with collaborative and individual **Activities**


relevant to the material presented in the text. These range from basic naked-eye and telescopic observing projects to opinion polls, surveys, group discussions, and astronomical research on the Web.

Chapter Review Summaries The Chapter Review Summaries, a primary review tool, are linked to the Learning Outcomes at the beginning of each chapter. Key Terms introduced in each chapter are listed again, in context and in boldface, along with key figures and page references to the text discussion.


Chapter Review

SUMMARY

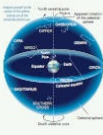
1 The **universe** (p. 7) is the totality of all space, time, matter, and energy. **Astronomy** (p. 7) is the study of the universe. In order of increasing size, the basic constituents of the cosmos are planets, stars, galaxies, galaxy clusters, and the universe itself. They differ enormously in scale—a factor of a billion billion from planet Earth to the entire observable universe.




2 The **scientific method** (p. 8) is a methodical approach employed by scientists to explore the universe around us in an objective manner. A **theory** (p. 8) is a framework of ideas and assumptions used to explain some set of observations and construct **theoretical models** (p. 8) that make predictions about the real world. These predictions in turn are amenable to further observational testing. In this way, the theory expands and science advances.



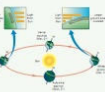
3 Early observers grouped the thousands of stars visible to the naked eye into patterns called **constellations** (p. 10), which they imagined were attached to a vast **celestial sphere** (p. 11) centered on Earth. Constellations have no physical significance, but are still used to label regions of the sky. The points where Earth's axis of rotation intersects the celestial sphere are called the north and south **celestial poles** (p. 12). The line where Earth's equatorial plane cuts the celestial sphere is the **celestial equator** (p. 12).



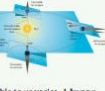
4 The nightly motion of the stars across the sky is the result of Earth's **rotation** (p. 12) on its axis. The time from one noon to the next is called a **solar day** (p. 13). The time between successive risings of any given star is a **sidereal day** (p. 13). Because of Earth's **revolution** (p. 14) around the Sun, we see different stars at night at different times of the year, and the Sun appears to move relative to the stars. The Sun's apparent yearly path around the celestial sphere (or the plane of Earth's orbit around the Sun) is called the **ecliptic** (p. 14).



5 We experience **seasons** (p. 16) because Earth's rotation axis is inclined to the ecliptic plane. At the **summer solstice** (p. 15), the Sun is highest in the sky and the length of the day is greatest. At the **winter solstice** (p. 16), the Sun is lowest and the day is shortest. At the **vernal** (p. 17) and **autumnal equinoxes** (p. 17), Earth's axis of rotation is perpendicular to the line joining Earth to the Sun, so day and night are of equal length. Because of **precession** (p. 17), the slow "wobble" of Earth's axis due to the influence of the Moon, the orientation of Earth's axis changes slowly over time. As a result, the particular constellations visible during any given season change over the course of thousands of years.



6 The Moon emits no light of its own, but instead shines by reflected sunlight. As the Moon orbits Earth, we see lunar **phases** (p. 18) as the amount of the Moon's sunlit face visible to us varies. A **lunar eclipse** (p. 20) occurs when the Moon enters Earth's shadow.



Instructor Resources



MasteringAstronomy® is the most widely used and most advanced astronomy tutorial and assessment system in the world. By capturing the step-by-step work of students nationally, MasteringAstronomy® has established an unparalleled database of learning challenges and patterns. Using this student data, a team of renowned astronomy education researchers has refined every activity and problem. The result is a library of activities of unique educational effectiveness and assessment accuracy. MasteringAstronomy® provides students with two learning systems in one: a dynamic self-study area and the ability to participate in online assignments.

MasteringAstronomy® provides instructors with a fast and effective way to assign uncompromising, wide-ranging online homework assignments of just the right difficulty and duration. The tutorials coach 90 percent of students to the correct answer with specific wrong-answer feedback. Powerful post-diagnostics allow instructors to assess the progress of their class as a whole or to quickly identify an individual student's areas of difficulty. Tutorials built

around text content and all the end-of-chapter problems from the text are available in MasteringAstronomy®. A media-rich self-study area is included that students can use whether the instructor assigns homework or not.

Instructor Guide Revised by James Heath (Austin Community College), this online guide provides sample syllabi and course schedules; an overview of each chapter; pedagogical tips; useful analogies; suggestions for classroom demonstrations; writing questions, selected readings, and answers/solutions to the end-of-chapter Review and Discussion Questions and Problems; and additional references and resources. ISBN 0-13-455418-3

Test Bank An extensive file of approximately 2800 test questions, newly compiled and revised for the ninth edition. The questions are organized and referenced by chapter section and by question type. The ninth edition Test Bank has been thoroughly revised and includes many new Multiple-Choice and Essay questions for added conceptual emphasis. This Test Bank is available in both Microsoft® Word and TestGen® formats. ISBN 0-13-455413-2

Instructor Resource Area in Mastering Astronomy This instructor resource area resides in MasteringAstronomy® and provides every electronic asset professors will need in and out of the classroom. The area not only contains an Instructor's Resource Manual, but also all text figures in jpeg and PowerPoint formats, including additional images, star charts, as well as the animations and videos from the MasteringAstronomy® Study Area. The area also contains TestGen®, an easy-to-use, fully networkable program for creating tests ranging from short quizzes to long exams. Questions from the Test Bank are supplied, and professors can use the Question Editor to modify existing questions or create new questions. It also contains chapter-by-chapter lecture outlines and conceptual "clicker" questions in PowerPoint. It is available in both PC and Mac formats.

Instructor Resource Center The Pearson Instructor Resource Center contains everything found on the Instructor Resource Area in MasteringAstronomy®. It provides virtually every electronic asset professors will need in and out of the classroom. This includes all text figures in jpeg and PowerPoint formats, as well as the animations and videos from the Mastering Astronomy® Study Area. It also contains TestGen®, an easy-to-use, fully networkable program for creating tests ranging from short quizzes to long exams. Questions from the Test Bank are supplied, and professors can use the Question Editor to modify existing questions or create new questions. This also contains chapter-by-chapter lecture outlines in PowerPoint and conceptual "clicker" questions in PowerPoint. ISBN 0-13-460971-9

Student Resources



This homework, tutorial, and assessment system is uniquely able to tutor each student individually by providing students with instantaneous feedback specific to their wrong answers, simpler subproblems upon request when they get stuck, and partial credit for their method(s) used. Students also have access to a self-study area that contains practice quizzes, self-guided tutorials, narrated and interactive figures, animations, videos, and more.

Pearson eText 2.0 is available through MasteringAstronomy®, either automatically when MasteringAstronomy® is packaged with new books, or available as a purchased upgrade online. Allowing the students to access the text wherever they have access to the Internet, Pearson eText comprises the full text, including figures that can be enlarged for better viewing, and embedded narrated and interactive figures where relevant. Within Pearson eText students are also able to pop up definitions and terms to help with vocabulary and the reading of the material. Students can also take notes in Pearson eText using the annotation feature.

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Sky and Telescope Based on the most popular amateur astronomy magazine, this special student supplement contains nine articles by Evan Skillman, each with a general overview and four question sets focused on the issues professors most want to address in this course: General Review, Process of Science, Scale of the Universe, and Our Place in the Universe. ISBN 0-321-70620-X

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Acknowledgments

Throughout the many drafts that have led to this book, we have relied on the critical analysis of many colleagues. Their suggestions ranged from the macroscopic issue of the

book's overall organization to the minutiae of the technical accuracy of each and every sentence. We have also benefited from much good advice and feedback from users of the first eight editions of the text. To these many helpful colleagues, we offer our sincerest thanks.

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The publishing team at Pearson has assisted us at every step along the way in creating this text. Special thanks go to content producers Alyc Helms, Lizette Faraji, and Rebecca Groves, who managed the many conflicting variables and looming deadlines that are a part of a multifaceted publication such as this. Executive editor Nancy Whilton steered this edition through all its phases. Production manager Jason Hammond of SPi Global has done an excellent job of tying together the threads of this very complex project, made all the more complex by the necessity of combining text, art, and electronic media into a coherent whole.

Special thanks are in order to cover and interior designer Jeff Puda for making the ninth edition look spectacular and to Marilyn Perry and Mark Ong for guiding the overall look of the book. We would also like to express our appreciation to Jenny Moryan for updating and maintaining the media resources in the MasteringAstronomy® Study Area and to Christina Cavalli, author of the MasteringAstronomy® Narrated Figures.

Eric Chaisson
Steve McMillan


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ASTRONOMY TODAY^{9e}

Based on cutting-edge scientific and education research, authors bring together the story of astronomy

Exoplanets 15

Planetary Systems Beyond Our Own



This artist's conception depicts the planet Kepler-90c, a super-Earth exoplanet (about 1.5 times Earth's radius) orbiting a star similar to the Sun in mass, temperature, and age. The exoplanet is in the habitable zone, the region around the star where the equilibrium temperature of the planet allows liquid water to exist on its surface. Without knowing its mass, composition, or atmosphere, however, scientists cannot yet be certain whether the planet is capable of harboring life. (NASA/JPL-Caltech)

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The Big Picture

There are currently thousands of known exoplanets, or planets orbiting stars other than the Sun, and likely billions exist in the Milky Way Galaxy alone. Until 1995, however, our solar system was the only planetary system upon which to build planetary formation theories. Current research involves discovering and characterizing individual planetary systems, describing the population as a whole, and improving theories of planetary system formation. With this work, astronomers seek to answer two fundamental questions: Is Earth unique? and Are we alone in the Universe? Just as comparative planetology of our eight neighboring planets guides our knowledge of Earth's history, extrasolar planets have much to teach us about how they all came to be in the first place.

SECTION 15.5 Our Solar System Unravelled 879

FIGURE 15.14 Whole New Worlds This artist's illustration depicts the variety of planets discovered by the Kepler telescope—some rocky planets, others perhaps with rocky or water-rich surfaces. Additional observations will be needed to pin down the nature of these alien worlds. (NASA/JPL-Caltech)

At the time of writing, more than 50 confirmed super-Earths and seven exo-Earths are known to orbit in or near the habitable zones of their parent stars. They are indicated as labeled points on Figure 15.13. A few of these planets with measured radii also have estimated masses, implying densities consistent with rocky/metallic terrestrial composition. The other planets in Figure 15.13 are of unknown density and hence composition. More of the planets in the "outer" edge of the habitable zone than on the "cool" side, but this is just another aspect of the observational bias described earlier—planets in close orbits are the most likely to be detected. Figure 15.14 places some of the new findings into perspective, again with a considerable amount of artistic license.

Many planet hunters are confident that within the next decade (or sooner), observational techniques will reach the level of sophistication at which jovian and even terrestrial planets similar to those in our solar system should be readily detectable—if they exist. Advances during the next decade will either bring numerous detections of extrasolar planets in "solar system" orbits or allow astronomers to conclude that systems like our own really are a small minority. Either way, the consequences are profound.

CONCEPT CHECK

How does the condensation theory accommodate the dissimilarities between the properties of planets in our solar system and those of the known extrasolar planets?

CONCEPT CHECK

Why are astronomers so interested in Earth-crossing asteroids?

CONCEPT CHECK

How does the condensation theory accommodate the dissimilarities between the properties of planets in our solar system and those of the known extrasolar planets?

The Big Question

People from all walks of life—not just astronomers—eagerly await the discovery of a true Earth-like planet orbiting another star. When will Earth's twin be found, and will it have solid surface, water oceans, and a hospitable atmosphere? Most intriguing of all, will it be inhabited? We live at a remarkable time when we are actually addressing—and often answering—some of the most profound questions that human beings have pondered for thousands of years.

NEW! Data Points side bars, based on data captured from thousands of students using MasteringAstronomy®, alert students to the most common mistakes made in working problems on related topics.

Discussion

MasteringAstronomy® provides a variety of resources to help you understand and master the concepts in this chapter. Visit the MasteringAstronomy Study Area for quizzes, animations, videos, interactive figures, and self-guided tutorials.

Conceptual Self-Test: Multiple Choice

Problems

Activities

Significantly Enhanced!

Big Picture theme, now woven throughout each chapter, helps students see connections between topics in the text and their everyday lives. Each chapter begins with a **Big Picture** opener, continues with **Big Picture** callouts throughout the chapter, and concludes with end-of-chapter **Big Picture** Questions, engaging students in a titillating narrative and helping them recognize themes.

340 CHAPTER 14 Solar System Debris

DATA POINTS

Comets

More than 60 percent of students had difficulty connecting the physical properties of comets to what we actually observe. Some points to remember:

- Comets originate in the outer solar system, in the Kuiper belt and Oort cloud, as small, icy bodies. Only the few comets with highly eccentric orbits that bring them near the Sun ever become visible from Earth.
- As a comet approaches the Sun, it warms up and its icy surface begins to vaporize. Comets "blow" when sunlight reflects off the gas and dust released by the Sun's heat, making them visible from Earth.
- Unlike meteors and other transient phenomena, comets move slowly across the sky relative to the stars over a period of weeks or even months.
- Comets tails don't follow behind the comet as it moves, but rather are always directed away from the Sun, blown outward by the solar wind.

14.2 Comets

Comets are usually discovered as faint, fuzzy patches of light on the sky while they are still several astronomical units away from the Sun. Traveling in a highly elliptical orbit with the Sun at one focus (Figure 14.2), a comet brightens and develops an extended tail as it nears the Sun. (The name "comet" derives from the Greek word *κόμη*, meaning "hair.") As the comet departs the Sun's vicinity, its brightness and tail diminish until it once again becomes a faint point of light receding into the distance. Like the planets, comets emit no visible light of their own—they shine by reflected (or remitted) sunlight. Each year, a few dozen are detected as they pass through the inner solar system. Many more must pass by unseen.

Comet Appearance and Structure

The various parts of a typical comet are shown in Figure 14.8. Even through a large telescope, the nucleus, or main solid part, of a comet is no more than a minute point of light. A typical cometary nucleus is extremely small—only a few kilometers in diameter. During most of the comet's orbit, far from the Sun, only this frozen nucleus exists. When a comet comes within a few astronomical units of the Sun, however, its icy surface becomes too warm to remain stable. Part of the comet becomes gaseous and expands into space, forming a diffuse coma ("halo") of dust and evaporated gas around the nucleus. The process by which a solid ice in this case ice changes directly into a gas, without first becoming liquid, is called **sublimation**. Frozen carbon dioxide (dry ice) on Earth provides a familiar example of this process. In space, sublimation is the rule, rather than the exception, for the behavior of ice when it is exposed to heat.

The coma becomes larger and brighter as the comet nears the Sun. At maximum size, the coma can measure 100,000 km in diameter—almost as large as Saturn or Jupiter. Enveloping the coma, an invisible **hydrogen envelope**, usually distorted by the solar wind, stretches across millions of kilometers of space. The comet's tail, which is most pronounced when the comet is closest to the Sun and the rate of sublimation from the nucleus is greatest, is much larger still, sometimes spanning as much as 1 AU.

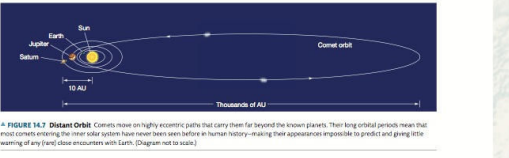


FIGURE 14.7 Distant Orbit Comets move on highly eccentric paths that carry them far beyond the known planets. Their long orbital periods mean that most comets entering the inner solar system have never been seen before in human history—making their appearances impossible to predict and giving little warning of any (very) close encounters with Earth. (Diagram not to scale)

UPDATED! New author-team member, Emily Rice, applies education research to restructure end-of-chapter content in the ninth edition. Newly redesigned discussion sections as well as **Individual and Collaborative Activities** now facilitate an active classroom.

New discoveries engage students by helping them realize astronomy is happening right now and all around them

Mars 10



A Near Miss for Life?

This 2015 "selfie" shows NASA's Curiosity rover atop a small hill in the floor of Gale crater on Mars. The region was probably once filled with water, making it a prime target for Curiosity's onboard chemistry lab. Who took the photograph? The rover's supporting arm contained the camera, which was removed from the photo. The white dot in the foreground marks one of several holes made in the Martian bedrock by Curiosity's drill. (NASA/JPL-Caltech/MSSS)

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The Big Picture

Mars has been reconnoitered by humans more than any other cosmic object beyond Earth and the Moon. It intrigues us, frustrates us, and tempts us to visit. Although the surface of Mars today is as dry as any desert on Earth, many scientists think it was much wetter billions of years ago, when its atmosphere was thicker and the climate was warmer. Mars remains high on most astronomers' list of likely sites for life in the solar system.

Chapters 10–12 include late-breaking updates from the Juno and Mars Maven Missions including discoveries from the Curiosity and new information on Saturn.

Chapters 14 & 15 reflect recent discoveries regarding solar system debris from the Dawn Mission at Ceres, the New Horizon Mission, and exoplanet searches and properties including the recent discovery of a potentially habitable planet around Proxima Centaur.

Solar System Debris 14



Keys to Our Origin

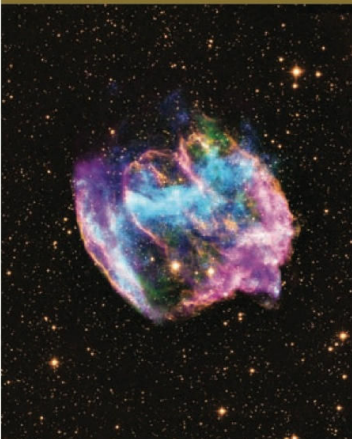
Pluto is the largest "small" body in the solar system. For decades it hid its secrets from even the largest telescopes, appearing as little more than a point of light in images taken from afar and forcing scientists to rely on models and speculation in their efforts to understand its nature. In 2015, NASA's New Horizons mission flew by the dwarf planet, capturing this image from a distance of 14,000 km. The probe's images have revealed astonishing and completely unexpected surface and atmospheric features, opening the conventional wisdom about the outer solar system. (NASA)

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The Big Picture

Hundreds of thousands of other celestial bodies—asteroids, comets, Kuiper-belt objects, and meteoroids—are known to orbit the Sun in addition to the eight planets and their moons. Astronomers estimate that there are more than a billion more minor bodies still to be discovered. These objects are small and of negligible mass compared with the planets and their moons, but more than the planets themselves, they hold a vital record of the formative stages of our planetary system.

Neutron Stars and Black Holes 22



Strange States of Matter

This striking image, a composite of individual X-ray (blue and green), infrared (yellow), and radio (red) images, shows SN1987A, a violent supernova event that marked the death of a massive star in our Galaxy about 100,000 years ago. Its distorted, asymmetric shape and unusual composition suggest to many astronomers that SN1987A may actually be a rare example of a hypernova—the most energetic class of stellar explosion known—with a black hole at its center. (JPL/NASA/CXC/MOS. Lopez et al.; infrared: Patmor; Radio: NSF/NRAO/VLA)

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The Big Picture

The almost unimaginable violence of supernova explosions may create objects so extreme in their behavior that they require us to reconsider some of our most cherished laws of physics. They open up a science fiction writer's dream of fantastic phenomena that border on reality. They may even one day force scientists to construct a whole new theory of the universe.

Chapter 22 now contains the latest information on gravitational radiation and black holes, particularly the connection between hypernovae and black holes.